Set 2: Cosmic Geometry

Newton vs Einstein

- Even though locally Newtonian gravity is an excellent approximation to General Relativity, in cosmology we deal with spatial and temporal scales across which the global picture benefits from a basic understanding of General Relativity
- An example is: as we have seen in the previous set of notes, it is much more convenient to think of the space between galaxies expanding rather than galaxies receding through space
- While latter is a good description locally, its preferred coordinates place us at center and does not allow us to talk about distances beyond which galaxies are receding faster than light though these distances as we shall see are also not directly observable
- To get a global picture of the expansion of the universe we need to think geometrically, like Einstein not Newton
- Best of both: think globally (Einstein), act locally (Newton)

Gravity as Geometry

- Einstein says Gravity as a force is really the geometry of spacetime
- Force between objects is a fiction of geometry imagine the curved space of the 2-sphere e.g. the surface of the earth
- Two people walk from equator to pole on lines of constant longitude
- Intersect at poles as if an attractive force exists between them
- Both walk on geodesics or straight lines of the shortest distance



Gravity as Geometry

- General relativity has two aspects
 - A metric theory: geometry tells matter how to move
 - Field equations: matter tells geometry how to curve
- Metric defines distances or separations in the spacetime and freely falling matter moves on a path that extremizes the distance
- Expansion of the universe carries two corresponding pieces

 Friedmann-Robertson-Walker geometry or metric tells matter, including light, how to move – allows us to chart out the expansion with light

 Matter content of the universe tells it how to expand – allows us to infer the components of the universe

• Useful to separate out these two pieces both conceptually and for understanding alternate cosmologies

- FRW geometry = homogeneous and isotropic on large scales
- Universe observed to be nearly isotropic (e.g. CMB, radio point sources, galaxy surveys)
- Copernican principle: we're not special, must be isotropic to all observers (all locations)

Implies homogeneity

Verified through galaxy redshift surveys

• FRW cosmology (homogeneity, isotropy & field equations) generically implies the expansion of the universe, except for special unstable cases

Isotropy & Homogeneity

- Isotropy: CMB isotropic to 10^{-3} , 10^{-5} if dipole subtracted
- Redshift surveys show return to homogeneity on the >100Mpc scale



• Spatial geometry is that of $\Omega_0 > 1$ constant curvature Positive: sphere $\Omega_0 < 1$ Negative: saddle Flat: plane • Metric tells us how to measure distances $\Omega_0 = 1$ on this surface MAP990006

- Closed: sphere of radius R and (real) curvature $K = 1/R^2$
- Suppress 1 dimension α represents total angular separation between two points on the sky (θ₁, φ₁) and (θ₂, φ₂)
- Embedding diagram (to help visualization, universe is on surface):



- Geometry tells matter how to move: take (null) geodesic motion for light along this generalized sense of longitude or radial distance D
- This arc distance is the distance our photon traveler experiences
- We receive light from two different trajectories as observer at pole
- Compared with our Euclidean expectation that the angle between the rays should be related to the separation at emission Σ as dα ≈ Σ/D the angular size appears larger because of the "lensing" magnification of the background
- This leads to the so called angular diameter distance the most relevant sense of distance for the observer
- In General Relativity, we are free to use any distance coordinate we like but the two have distinct uses

• To define the angular diameter distance, look for a D_A such that

 $d\Sigma = D_A d\alpha$

Draw a circle at the distance D, its radius is $D_A = R \sin(D/R)$



- Angular diameter distance
- Positively curved geometry $D_A < D$ and objects are further than they appear
- Negatively curved universe R is imaginary and R sin(D/R) = i|R| sin(D/i|R|) = |R| sinh(D/|R|)
 and D_A > D objects are closer than they appear
- Flat universe, $R \to \infty$ and $D_A = D$

- Now add that point 2 may have a different radial distance
- What is the distance dΣ between x₁ = (θ₁, φ₁, D₁) and point 2
 x₂ = (θ₂, φ₂, D₂), separated by dα in angle and dD in distance?



Angular Diameter Distance

• For small angular and radial separations, space is nearly flat so that the Pythagorean theorem holds for differentials

$$d\Sigma^2 = dD^2 + D_A^2 d\alpha^2$$

• Now restore the fact that the angular separation can involve two angles on the sky - the curved sky is just a copy of the spherical geometry with unit radius that we were suppressing before

$$d\Sigma^2 = dD^2 + D_A^2 d\alpha^2$$
$$= dD^2 + D_A^2 (d\theta^2 + \sin^2\theta d\phi^2)$$

- D_A useful for describing observables (flux, angular positions)
- *D* useful for theoretical constructs (causality, relationship to temporal evolution)

Alternate Notation

• Aside: line element is often also written using D_A as the coordinate distance

$$dD_{A}^{2} = \left(\frac{dD_{A}}{dD}\right)^{2} dD^{2}$$
$$\left(\frac{dD_{A}}{dD}\right)^{2} = \cos^{2}(D/R) = 1 - \sin^{2}(D/R) = 1 - (D_{A}/R)^{2}$$
$$dD^{2} = \frac{1}{1 - (D_{A}^{2}/R)^{2}} dD_{A}^{2}$$

and defining the curvature $K = 1/R^2$ the line element becomes

$$d\Sigma^{2} = \frac{1}{1 - D_{A}^{2}K} dD_{A}^{2} + D_{A}^{2} (d\theta^{2} + \sin^{2}\theta d\phi^{2})$$

where K < 0 for a negatively curved space

Line Element or Metric Uses

• Metric also defines the volume element

$$dV = (dD)(D_A d\theta)(D_A \sin \theta d\phi)$$
$$= D_A^2 dD d\Omega$$

where $d\Omega = \sin \theta d\theta d\phi$ is solid angle

- Most of classical cosmology boils down to these three quantities, (comoving) radial distance, (comoving) angular diameter distance, and volume element
- For example, distance to a high redshift supernova, angular size of the horizon at last scattering and BAO feature, number density of clusters...

Comoving Coordinates

- Remaining degree of freedom (preserving homogeneity and isotropy) is the temporal evolution of overall scale factor
- Relates the geometry (fixed by the radius of curvature *R*) to physical coordinates a function of time only

 $d\sigma^2 = a^2(t)d\Sigma^2$

our conventions are that the scale factor today $a(t_0) \equiv 1$

- Similarly physical distances are given by d(t) = a(t)D,
 d_A(t) = a(t)D_A.
- Distances in upper case are comoving; lower, physical Comoving coordinates do not change with time Simplest coordinates to work out geometrical effects

Time and Conformal Time

• Spacetime separation (with c = 1)

$$ds^{2} = -dt^{2} + d\sigma^{2}$$
$$= -dt^{2} + a^{2}(t)d\Sigma^{2}$$

• Taking out the scale factor in the time coordinate

$$ds^2 \equiv a^2(t) \left(-d\eta^2 + d\Sigma^2 \right)$$

 $d\eta = dt/a$ defines conformal time – useful in that photons travelling radially from observer on null geodesics $ds^2 = 0$

$$\Delta D = \Delta \eta = \int \frac{dt}{a}$$

so that time and distance may be interchanged

FRW Metric

• Aside for advanced students: Relationship between coordinate differentials and space-time separation defines the metric $g_{\mu\nu}$

$$ds^2 \equiv g_{\mu\nu}dx^{\mu}dx^{\nu} = a^2(\eta)(-d\eta^2 + d\Sigma^2)$$

Einstein summation - repeated lower-upper pairs summed

- Usually we will use comoving coordinates and conformal time as the x^μ unless otherwise specified – metric for other choices are related by a(t)
- Aside: scale factor plays the role of a conformal rescaling (which preserves spacetime "angles", i.e. light cone and causal structure hence conformal time)

Horizon

- Distance travelled by a photon in the whole lifetime of the universe defines the horizon
- Since ds = 0, the horizon is simply the elapsed conformal time

$$D_{\text{horizon}}(t) = \int_0^t \frac{dt'}{a} = \eta(t)$$

- Horizon always grows with time
- Always a point in time before which two observers separated by a distance D could not have been in causal contact
- Horizon problem: why is the universe homogeneous and isotropic on large scales especially for objects seen at early times, e.g.
 CMB, when horizon small

Special vs. General Relativity

- From our class perspective, the big advantage of comoving coordinates and conformal time is that we have largely reduced general relativity to special relativity
- In these coordinates, aside from the difference between D and D_A, we can think of photons propagating in flat spacetime
- Now let's relate this discussion to observables
- Rule of thumb to avoid dealing with the expansion directly:
 - Convert from physical quantities to conformal-comoving quantities at emission
 - In conformal-comoving coordinates, light propagates as usual
 - At reception a = 1, conformal-comoving coordinates are physical, so interpret as usual

Recession as Expansion

• Recession and redshifting as coordinate expansion

Redshift

- Wavelength of light "stretches" with the scale factor
- The physical wavelength λ_{emit} associated with an observed wavelength today λ_{obs} (or comoving=physical units today) is

$$\frac{\lambda_{\text{emit}}}{\lambda_{\text{obs}}} = \frac{a(t_{\text{emit}})}{a(t_{\text{obs}})} = a(t_{\text{emit}})$$



so that the redshift of spectral lines measures the scale factor of the universe at t, 1 + z = 1/a.

• Interpreting the redshift as a Doppler shift, objects recede in an expanding universe

Hubble Parameter

- Since a = 1/(1 + z) is a direct observable, useful to cast it as the independent variable
- Convert using the expansion rate or Hubble parameter

$$H(t) \equiv \frac{1}{a}\frac{da}{dt} = \frac{d\ln a}{dt}$$

fractional change in the scale factor per unit time - $\ln a = N$ is also known as the e-folds of the expansion

• Cosmic time becomes

$$t = \int dt = \int \frac{d\ln a}{H(a)}$$

• Conformal time becomes

$$\eta = \int \frac{dt}{a} = \int \frac{d\ln a}{aH(a)}$$

Distance-Redshift Relation

- Given atomically known rest wavelength λ_{emit} , redshift can be precisely measured from spectra
- Combined with a measure of distance, distance-redshift
 D(z) ≡ D(z(a)) can be inferred given that photons travel
 D = Δη this tells us how the scale factor of the universe evolves with time.
- Related to the expansion history as

$$D(a) = \int dD = \int_{a}^{1} \frac{d\ln a'}{a'H(a')}$$
$$[d\ln a' = -d\ln(1+z) = -a'dz]$$
$$D(z) = -\int_{z}^{0} \frac{dz'}{H(z')} = \int_{0}^{z} \frac{dz'}{H(z')}$$

Hubble Law

• Note limiting case is the Hubble law

$$\lim_{z \to 0} D(z) = z/H(z=0) \equiv z/H_0$$

independently of the geometry and expansion dynamics

• Hubble constant usually quoted as as dimensionless h

$$H_0 = 100 h \,\mathrm{km \, s^{-1} Mpc^{-1}}$$

- Observationally $h \sim 0.7$ (see below)
- With c = 1, H₀⁻¹ = 9.7778 (h⁻¹ Gyr) defines the time scale (Hubble time, ~ age of the universe)
- As well as $H_0^{-1} = 2997.9 (h^{-1} \text{ Mpc})$ a length scale (Hubble scale ~ Horizon scale)

Standard Ruler

- Standard Ruler: object of known physical size λ
- Let's apply our rule of thumb: at emission the comoving size is Λ :

 $\lambda = a(t)\Lambda$

Now everything about light is normal: the object of comoving size Λ subtends an observed angle α on the sky α

$$\alpha = \frac{\Lambda}{D_A(z)}$$

• This is the easiest way of thinking about it. But we could also define an effective physical distance $d_A(z)$ which corresponds to what we would infer in a non expanding spacetime

$$\alpha \equiv \frac{\lambda}{d_A(z)} = \frac{a\Lambda}{d_A(z)} = \frac{\Lambda}{D_A(z)} \to d_A(z) = aD_A(z) = \frac{D_A(z)}{1+z}$$

Standard Ruler

- Since D_A → D_A(D_{horizon}) whereas (1 + z) unbounded, angular size of a fixed physical scale at high redshift actually increases with radial distance
- Paradox: the further away something is, the bigger it appears

– Easily resolved by thinking about comoving coordinates - a fixed physical scale λ as the universe shrinks and $a \rightarrow 0$ will eventually encompass the whole observable universe out to the horizon in comoving coordinates so of course it subtends a big angle on the sky!

– But there are no such bound objects in the early universe there is no causal way such bigger-than-the-horizon objects could form

• Knowing λ or Λ and measuring α and z allows us to infer the comoving angular diameter distance $D_A(z)$

Standard Candle

- Standard Candle: objects of same luminosity L, measured flux F
- Apply rules again: at emission in conformal-comoving coordinates
 - -L is the energy per unit time at emission

– Since $E \propto \lambda^{-1}$ and comoving wavelength $\Lambda \propto \lambda/a$ so comoving energy $\mathcal{E} \propto \Lambda^{-1} \propto aE$

- Per unit time at emission $\Delta t = a \Delta \eta$ in conformal time
- So observed luminosity today is $\mathcal{L} = \mathcal{E}/\Delta \eta = a^2 L$
- All photons must pass through the sphere at a given distance, so the comoving surface area is $4\pi D_A^2$
- Put this together to the observed flux at a = 1

$$F = \frac{\mathcal{L}}{4\pi D_A^2} = \frac{L}{4\pi D_A^2} \frac{1}{(1+z)^2}$$

Notice the flux is diminished by two powers of (1 + z)

Luminosity Distance

• We can again define a physical "luminosity" distance that corresponds to our non-expanding spacetime intuition

$$F \equiv \frac{L}{4\pi d_L^2}$$

• So luminosity distance

$$d_L = (1+z)D_A = (1+z)^2 d_A$$

• As
$$z \to 0$$
, $d_L = d_A = D_A$

• But as $z \to \infty$, $d_L \gg d_A$ - key to understanding Olber's paradox

Olber's Paradox Redux

• Surface brightness - object of physical size λ

$$S = \frac{F}{\Delta \Omega} = \frac{L}{4\pi d_L^2} \frac{d_A^2}{\lambda^2}$$

• In a non-expanding geometry (regardless of curvature), surface brightness is conserved $d_A = d_L$

S = const.

- each site line in universe full of stars will eventually end on
 surface of star, night sky should be as bright as sun (not infinite)
- Age finite so even if stars exist in the early universe, not all site lines end on stars
- In an expanding universe

$$S \propto (1+z)^{-4}$$

Olber's Paradox Redux

- But even as age goes to infinity and the number of site lines goes to 100%, surface brightness of distant objects (of fixed physical size) goes to zero
 - Angular size increases
 - Redshift of "luminosity" i.e. energy and arrival time dilation
- In an expanding universe

$$S \propto (1+z)^{-4}$$

Measuring D(z)

• Astro units side: since flux ratios are very large in cosmology, its more useful to take the log

$$m_1 - m_2 = -2.5 \log_{10}(F_1/F_2)$$

related to d_L by definition by inverse square law

$$m_1 - m_2 = 5 \log_{10}[d_L(z_1)/d_L(z_2)]$$

• To quote in terms of a single object, introduce absolute magnitude as the magnitude that would be measured for the object at 10 pc

$$m - M = 5 \log_{10}[d_L(z)/10 \text{pc}]$$

Knowing absolute magnitude is the same as knowing the absolute distance, otherwise distances are relative

Measuring D(z)

 If absolute magnitude unknown, then both standard candles and standard rulers measure relative sizes and fluxes – ironically this means that measuring the change in *H* is easier than measuring *H*₀ – acceleration easier than rate!

For standard candle, e.g. compare magnitudes low z_0 to a high z object - using the Hubble law $d_L(z_0) = z_0/H_0$ we have

$$\Delta m = m_z - m_{z_0} = 5 \log_{10} \frac{d_L(z)}{d_L(z_0)} = 5 \log_{10} \frac{H_0 d_L(z)}{z_0}$$

Likewise for a standard ruler comparison at the two redshifts

$$\frac{d_A(z)}{d_A(z_0)} = \frac{H_0 d_A(z)}{z_0}$$

• Distances are measured in units of h^{-1} Mpc.

Measuring D(z)

- Since z is a direct observable, in both cases $H_0D_A(z)$ is the measured quantity
- We can relate that back to $H_0D(z)$ recalling that

 $H_0 D_A = H_0 R \sin(H_0 D / H_0 R)$

and use h^{-1} Mpc as the unit for all lengths – furthermore, local observations are at distances much smaller than R so $H_0D_A = H_0D$ is a good approximation

• Then since $D(z) = \int dz / H(z)$ we have

$$H_0 D(z) = \int dz \frac{H_0}{H(z)}$$

• Fundamentally our low to high z comparison tells us the change in expansion rate $H(z)/H_0$

Supernovae as Standard Candles

- Type 1A supernovae are white dwarfs that reach Chandrashekar mass where electron degeneracy pressure can no longer support the star, hence a very regular explosion
- Moreover, the scatter in absolute magnitude is correlated with the shape of the light curve - the rate of decline from peak light, empirical "Phillips relation"



• Higher ⁵⁶Ni, brighter SN, higher opacity, longer light curve duration

Beyond Hubble's Law

- Type 1A are therefore "standardizable" candles leading to a very low scatter $\delta m \sim 0.15$ and visible out to high redshift $z \sim 1$
- Two groups in 1999
 found that SN more distant at a given redshift than expected
- Cosmic acceleration



Acceleration of the Expansion

• Using SN as a relative indicator (independent of absolute magnitude), comparison of low and high z gives

$$H_0 D(z) = \int dz \frac{H_0}{H}$$

more distant implies that H(z) not increasing at expect rate with z

• Intuition tells us (FRW dynamics shows) ordinary matter decelerates expansion since gravity is attractive

Or equivalently that expansion rate decreases as it expands

• Ordinary expectation is that

$$H(z>0) > H_0$$

e.g. $\propto (1+z)^{3/2}$, so that the Hubble parameter is higher at high redshift

Acceleration of the Expansion

- More dim (more distant) supernovae indicate that H(z) grows slower and hence is more constant than expected
- Take the limiting case where H(z) is strictly constant (a.k.a. de Sitter expansion)

$$H = \frac{1}{a} \frac{da}{dt} = \text{const}$$
$$\frac{dH}{dt} = \frac{1}{a} \frac{d^2a}{dt^2} - H^2 = 0$$
$$\frac{1}{a} \frac{d^2a}{dt^2} = H^2 > 0$$

- Indicates that the expansion of the universe is accelerating
- Notice that this a purely geometric inference and does not yet say anything about what causes acceleration – topic of next set of lectures on cosmic dynamics

Hubble Constant

- Getting H_0 itself is harder since need to know absolute distance d_L to objects: $H_0 = z_0/d_L$
- Hubble actually inferred too large a Hubble constant of $H_0 \sim 500$ km/s/Mpc



- Miscalibration of the Cepheid distance scale absolute measurement hard, checkered history
- Took 70 years to settle on this value with a factor of 2 discrepancy persisting until late 1990's which is after the projects which discovered acceleration were conceived!
- H_0 now measured as 73.04 ± 1.04 km/s/Mpc by SH0ES calibrating off local, geometric absolute distances including AGN water maser

Hubble Constant History

- Difficult measurement since local galaxies have peculiar motions and so their velocity is not entirely due to the "Hubble flow"
- A "distance ladder" of cross calibrated measurements
- Primary distance indicators cepheids, novae planetary nebula, tip of red giant branch, or globular cluster luminosity function, AGN water maser
- Use more luminous secondary distance indications to go out in distance to Hubble flow

Tully-Fisher, fundamental plane, surface brightness fluctuations, Type 1A supernova

Modern Distance Ladder



Maser-Cepheid-SN Distance Ladder

- Water maser around AGN, gas in Keplerian orbit
- Measure proper motion, radial velocity, acceleration of orbit



 Method 1: radial velocity plus orbit infer tangential velocity = distance × angular proper motion

$$v_t = d_A (d\alpha/dt)$$

• Method 2: centripetal acceleration and radial velocity from line infer physical size

$$a = v^2/R, \qquad R = d_A \theta$$

Maser-Cepheid-SN Distance Ladder

- Calibrate Cepheid period-luminosity relation in same galaxy
- SH0ES project then calibrates SN distance in galaxies with Cepheids

Also: consistent with recent HST parallax determinations of 10 galactic Cepheids (8% distance each) with $\sim 20\%$ larger H_0 error bars - normal metalicity as opposed to LMC Cepheids.

- Measure SN at even larger distances out into the Hubble flow
- Riess et al, arXiv: 2112.04510, 73.04 ± 1.04 km/s/Mpc
- Currently, this differs from the CMB/BAO distance ladder $H_0 = 67.66 \pm 0.42$ working from high redshifts at $\sim 5\sigma$