## Astro 448: Problem Set 3 Due October 24

## **1** Problem 1: Green's Method

• (a) From the continuity and Euler equations of the joint photon-baryon system (see notes) show

$$\ddot{\Theta} + \frac{\dot{R}}{1+R}\dot{\Theta} + k^2 c_s^2 \Theta = F(\eta) \tag{1}$$

$$F(\eta) = -\ddot{\Phi} - \frac{\dot{R}}{1+R}\dot{\Phi} - \frac{k^2}{3}\Psi$$
<sup>(2)</sup>

• (b) Take the solutions

$$\theta_a = (1+R)^{-1/4} \cos(ks) \tag{3}$$

$$\theta_b = (1+R)^{-1/4} \sin(ks) \tag{4}$$

and show that they solve the homogeneous F = 0 equation in the adiabatic approximation

• (c) Use the Greens method to construct the particular solution

$$\Theta(\eta) = C_1 \theta_a(\eta) + C_2 \theta_b(\eta) + \int_0^{\eta} d\eta' \frac{\theta_a(\eta')\theta_b(\eta) - \theta_a(\eta)\theta_b(\eta')}{\theta_a(\eta')\dot{\theta}_b(\eta') - \dot{\theta}_a(\eta')\theta_b(\eta')} F(\eta')$$
(5)

and give the expression in terms of  $\Theta(0)$ ,  $\dot{\Theta}(0)$ , R,  $\dot{R}(0)$ , s. Think of this as taking a set of impulsive forces on the oscillator and propagating their effect into a temperature perturbation at a later time.

- (d) Evalute the general solution for  $R, \Psi, \Phi$  all constant.
- (e) What you expect to happen qualitatively to the acoustic oscillations for initial conditions where there are no gravitational potentials initially and  $\Phi = \Psi$  only becomes substantial after horizon crossing  $k\eta = 1$  (take  $C_1 = C_2 = 0$ ). Remember that the Greens solution causally propagates an impulsive force. Argue that the appearance of a first peak that is consistent with adiabatic initial conditions is strong argument for inflation.