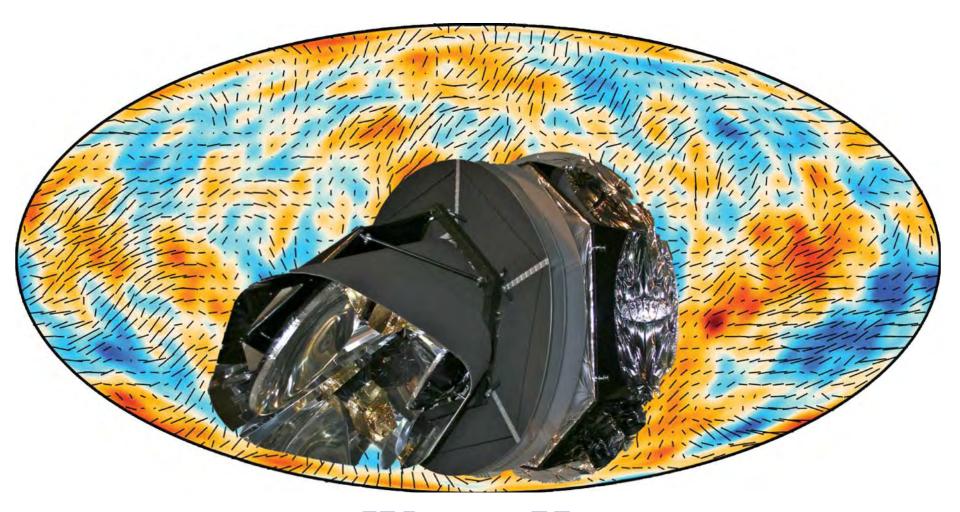
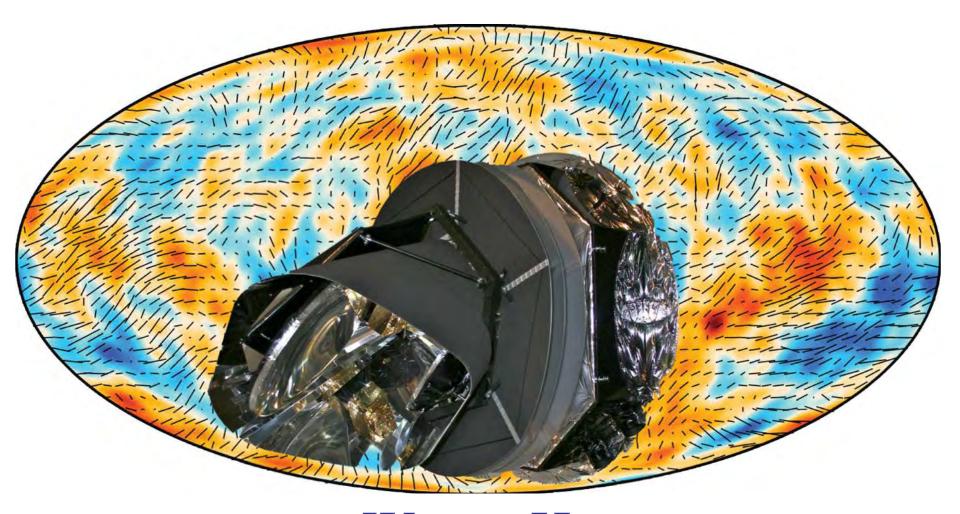
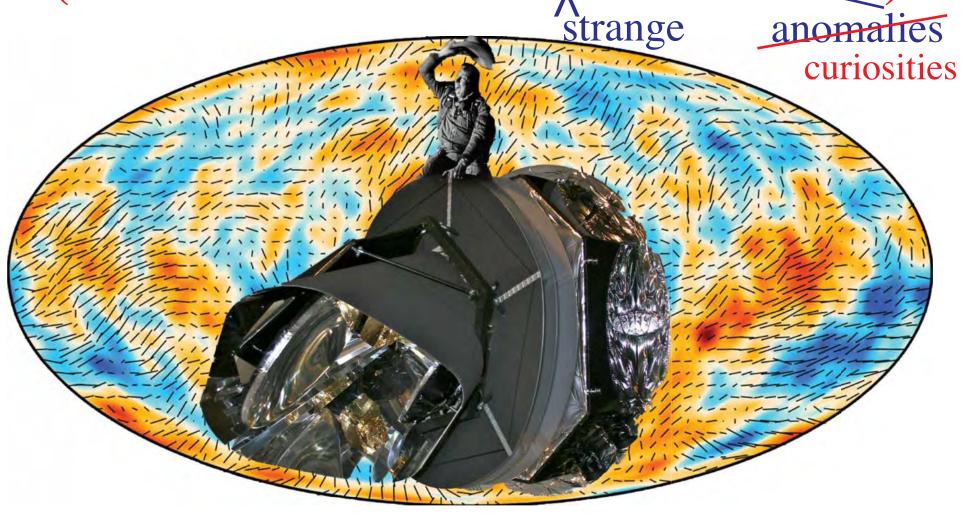
What we have learnt from Planck (organizer's title!)



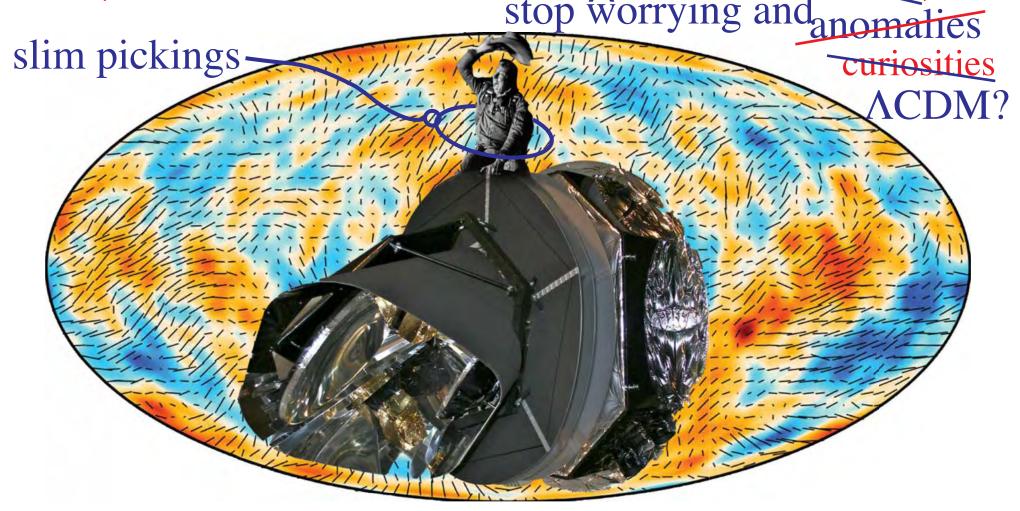
What we have learnt from Planck learned



What we have learnt from Planck (or how I learned to love tension) strange anomalies

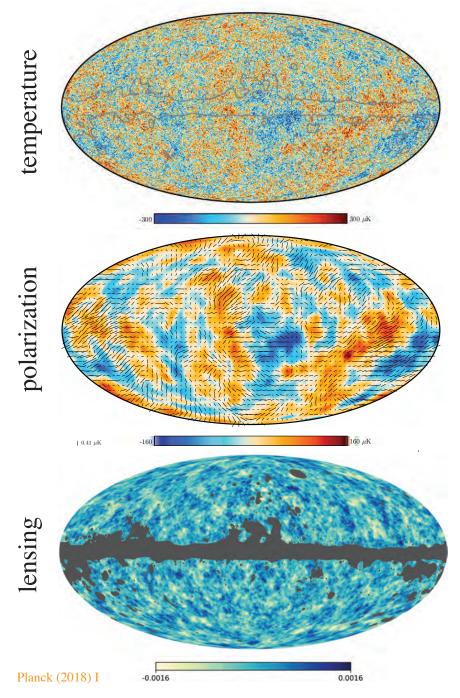


What we have learnt from Planck (or how I learned to love tension) stop worrying and anomalies



Paradigm of Precision Cosmology

- Precision measurements and maps
 - temperature
 - polarization
 - lensing
 - 9 frequencies
- Control over systematics
 - most recently polarization
- Accurate and precise theoretical predictions
 - Gaussian, adiabatic
 - $-\Lambda$ CDM



Near Perfection in 6 Numbers

 All this precision data described by
 6 ΛCDM parameters

 $-\Omega_c h^2$: CDM

 $-\Omega_b h^2$: baryons

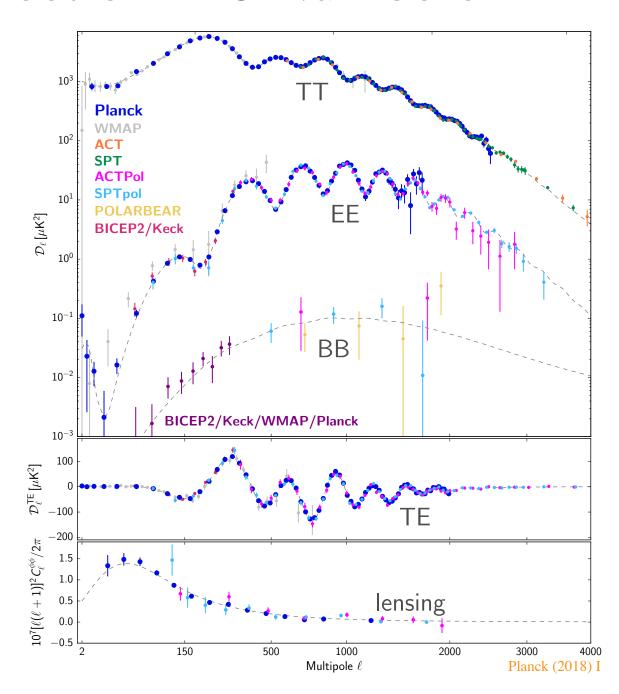
 $-\theta_s$: sound scale

 $-A_s$: amplitude

 $-n_s$: tilt

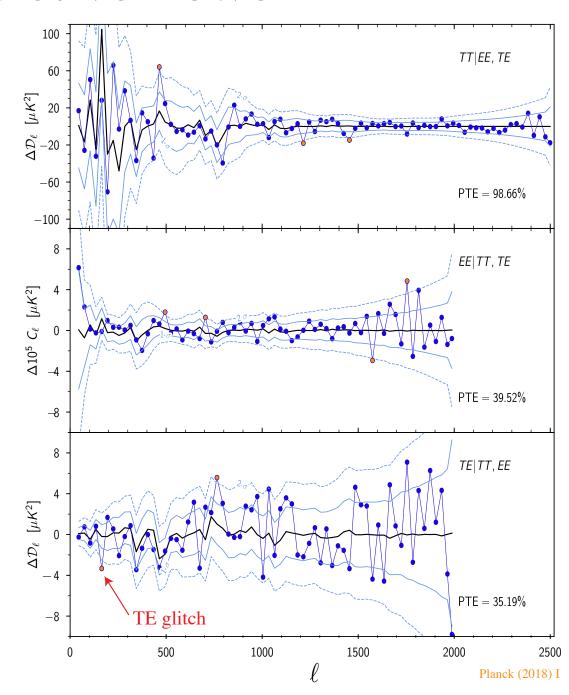
 $-\tau$: reionization

Measured
 to sub percent
 precision (except τ)



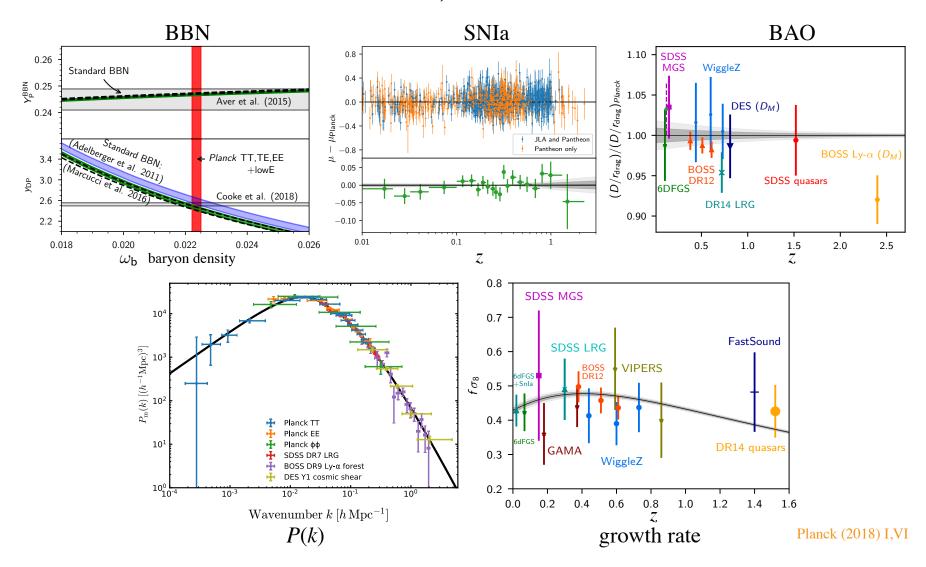
Predictive Power

- Small residuals from \Lambda CDM in various spectra
- Temp ↔ pol
 residuals in ΛCDM
 with reduced sample
 variance
- Largely consistent,
 but with high
 precision, moderately
 significant deviations
- $\sim 2\sigma$ outliers, expected but some also drive parameters



Predictive Power

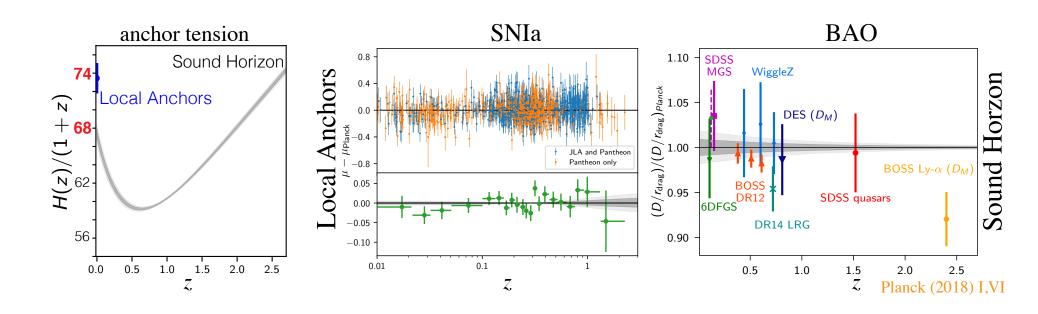
• Predicts all other observables, which direct measurements test



• Good agreement, even weak lensing, clusters, and yes H_0 (< 10%)

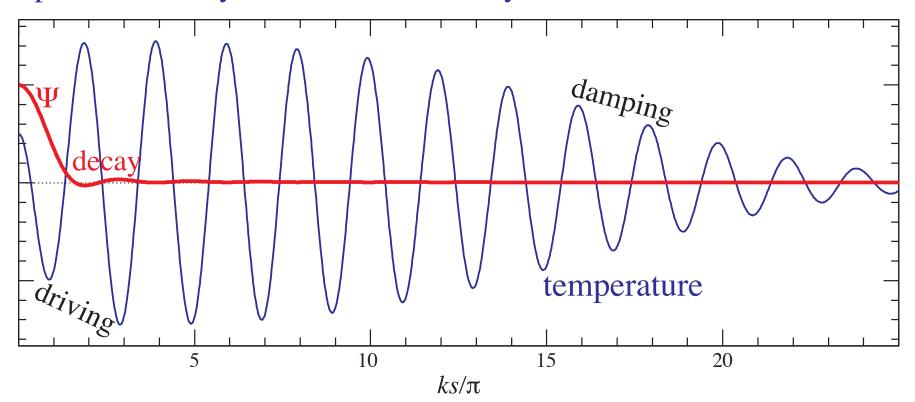
Anchors Sink \(\Lambda\)CDM?

- When distance ladder calibrated by CMB sound horizon, H_0 discrepant with local measurements at 4.4σ (Riess et al 2019)
- Relative distances forward/backwards by ladder: CMB to BAO to SN isolating discrepancy as anchors (e.g. Aylor et al. 2018)
- Relative distances $\sim \Lambda \text{CDM}$: little room for any new physics at intermediate redshifts to resolve



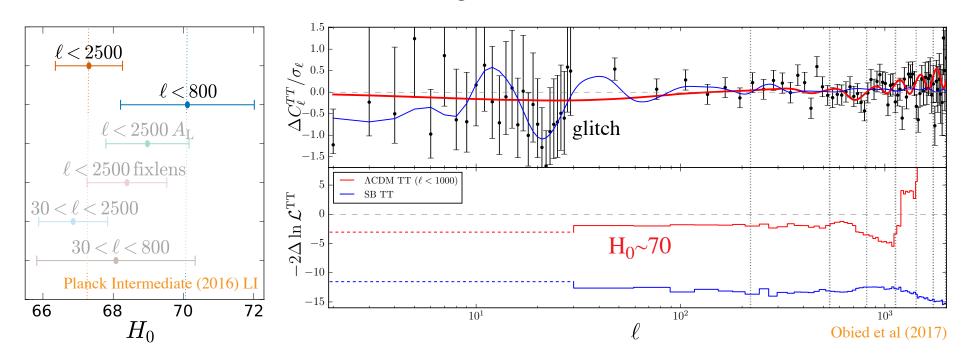
Driving in the Anchor

- CMB anchor is sound horizon, must calibrate propagation time
- $H(z < 10^3)$: with radiation, baryons fixed only $\Omega_c h^2$ unknown
- $\Omega_c h^2$ controls matter-radiation ratio and radiation driving from potential decay due to Jeans stability



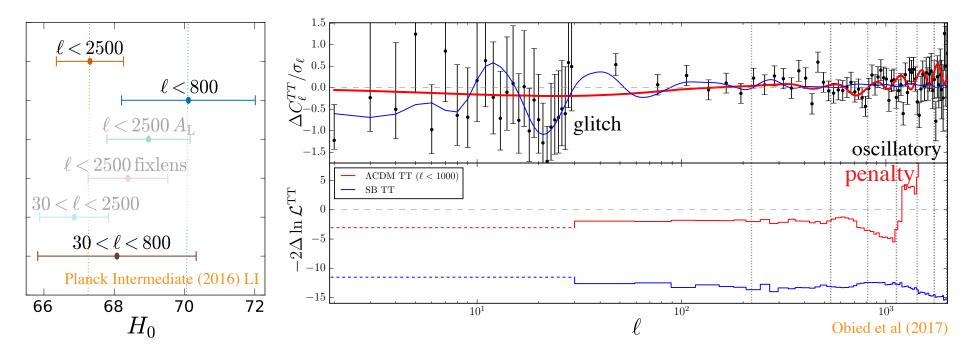
Glitch and Oscillatory Residuals

- Shifts in CMB anchor between low & high ℓ (Addison et al 2015)
- Low multipole glitch deficit of power looks like peaks should be higher, more driving, less matter, higher H_0
- High multipole oscillatory residuals: smoother acoustic peaks, less driving, more matter, lower H_0
 - also drives "lensing tension": Pavel Motloch's talk



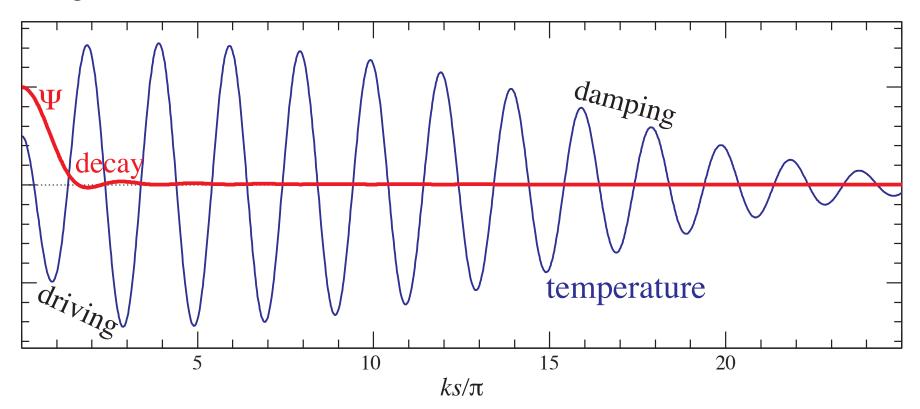
Glitch and Oscillatory Residuals

- Exclude low $\ell < 30$ glitch (Planck Intermediate 2016 LI)
- H_0 falls, consistent with high multipoles within errors
- Alternately marginalize over possible features during inflation, mild $\Delta \chi^2 \sim 12$ improvement, but 5 params
- Likewise H_0 returns to low value (Object at 2017)



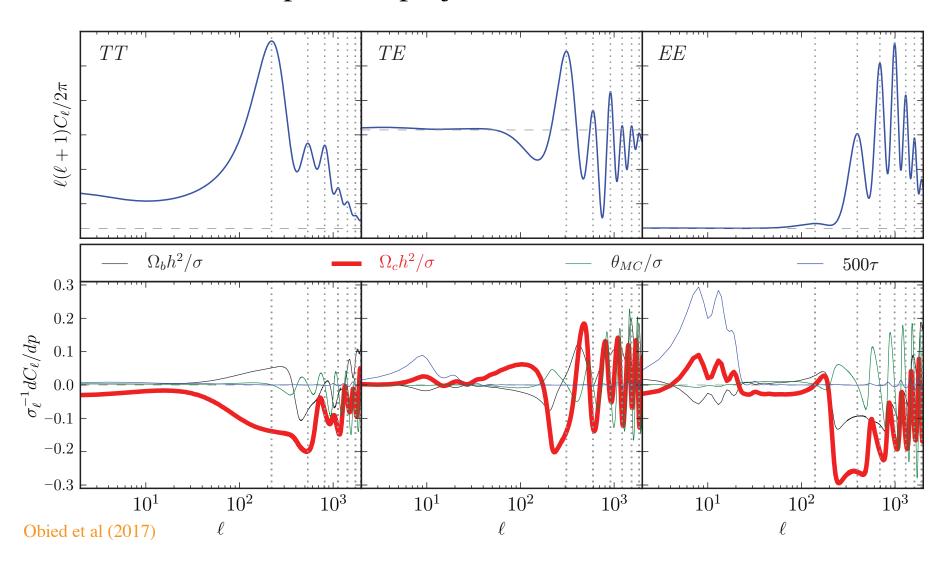
Driving and Oscillatory Residuals

- Oscillatory residuals indicate smoother peaks
- Driving sharpens the peaks
- Residuals indicate less driving, higher matter-radiation ratio
- Higher $\Omega_c h^2$, lower H_0



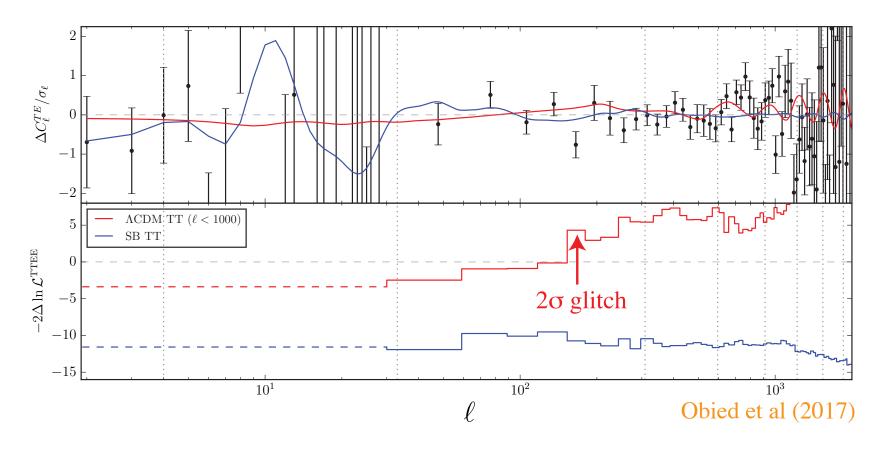
Driving and CDM

- Signatures of CDM $\Omega_c h^2$: TT amplitude and oscillatory residuals
- Polarization sharper test: projection effects for TT (Galli et al. 2014)



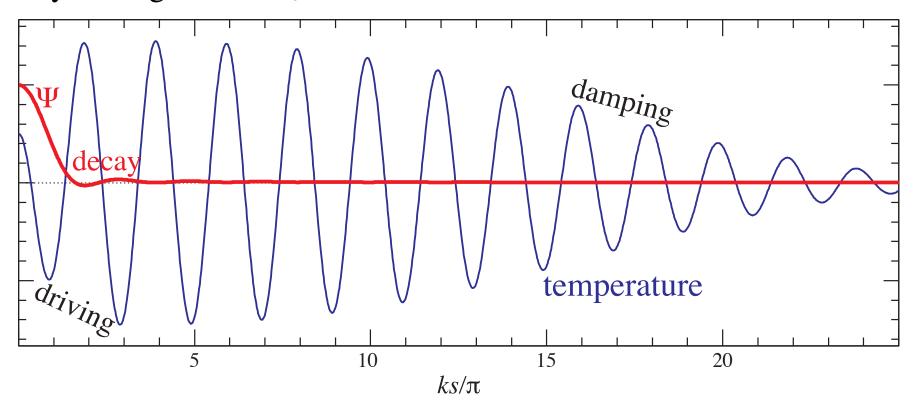
Polarization Signatures

- Polarization sensitivity provides independent calibration using $\ell < 1000$ TE Planck data
- TE glitch at $\ell \sim 165$ enhances sensitivity, since lowering $\Omega_c h^2$ (raising H_0) further raises predictions



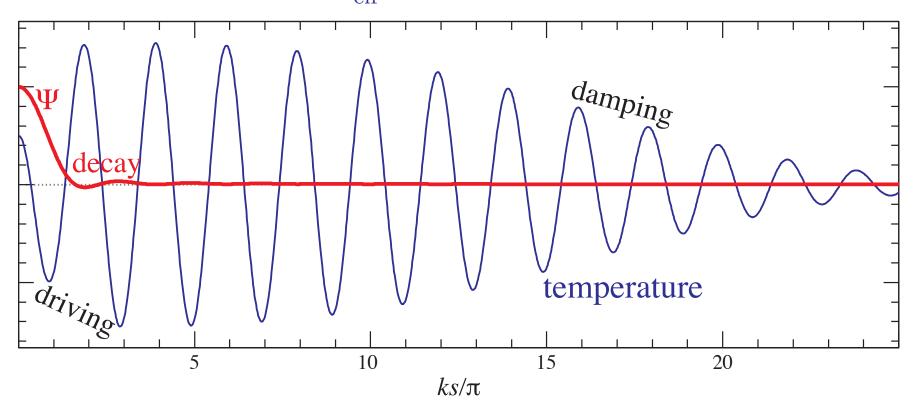
Dark Radiation

- Extra dark species whose energy density redshifts faster than matter change sound horizon calibration
- Raise $H(z < 10^3)$, lower sound horizon, raise H_0 at fixed θ_s
- Additional driving from Jeans stable $\Delta N_{\rm eff}$ radiation compensated by raising matter $\Omega_c h^2$



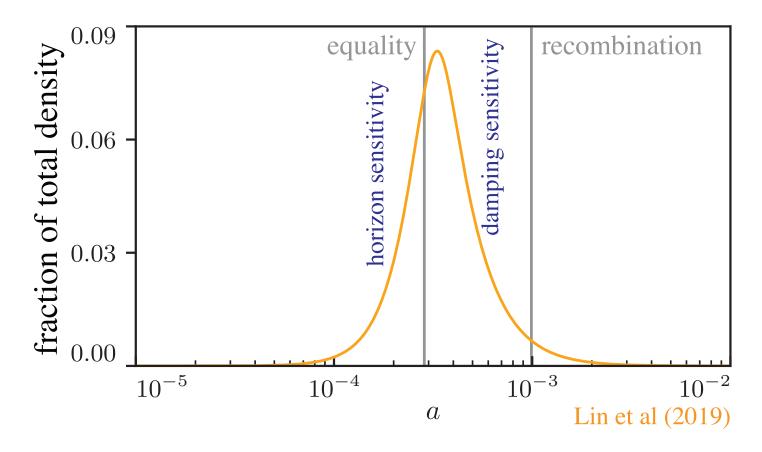
Driving and Damping

- But damping provides second standard ruler in diffusion scale
- Random walk distance of photon scales as harmonic mean between horizon and mean free path
- Consistency check that Λ CDM passes and constrains any additional radiation $\Delta N_{\rm eff}$



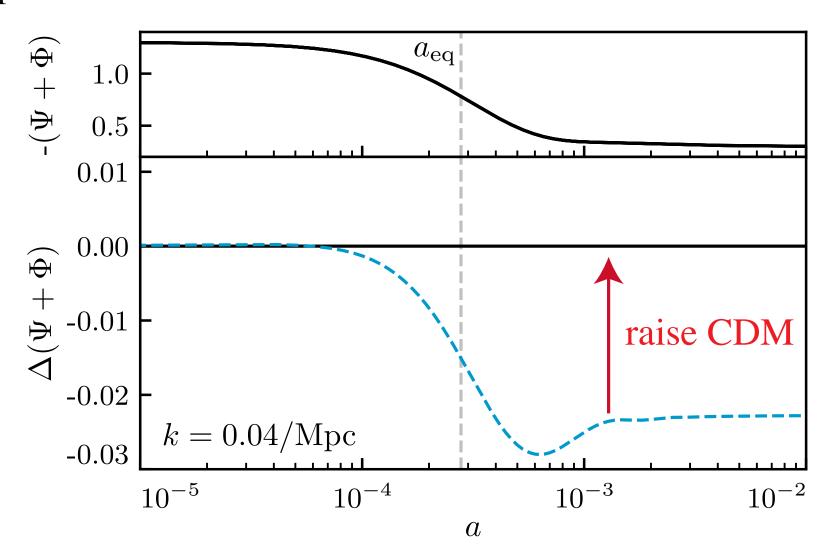
Dark Exotica

- Decreasing additional dark components during radiation domination changes damping vs sound horizon
- Reconcile ratio if timed exactly right

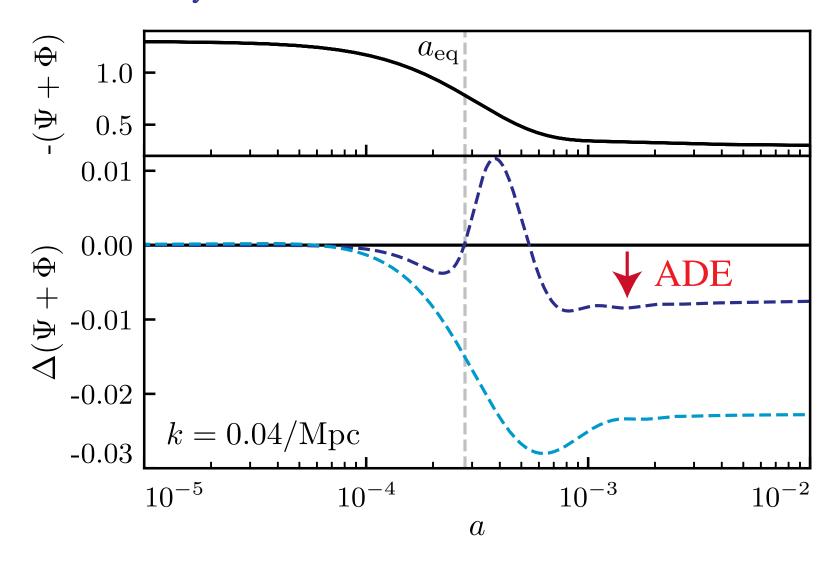


• Poulin et al (2018): specific anharmonic, periodic scalar potential

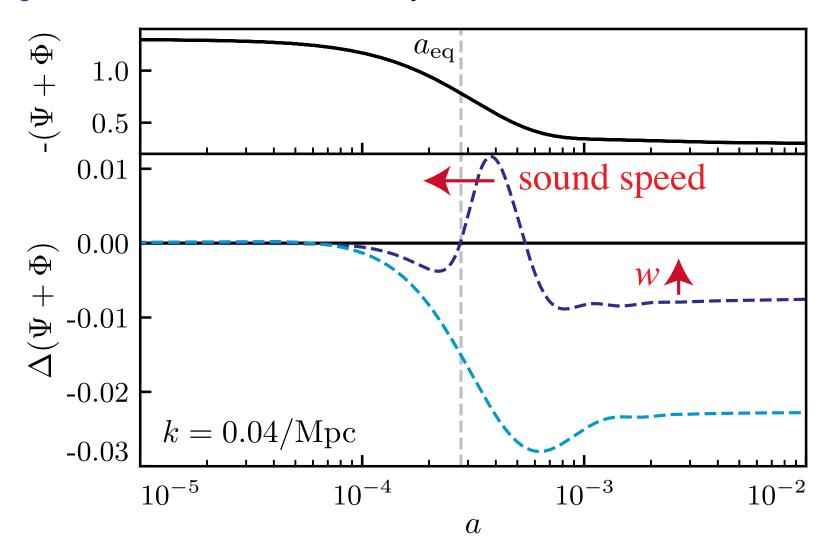
• Must compensate effect of raising $\Omega_c h^2$ which reduces decay of potential



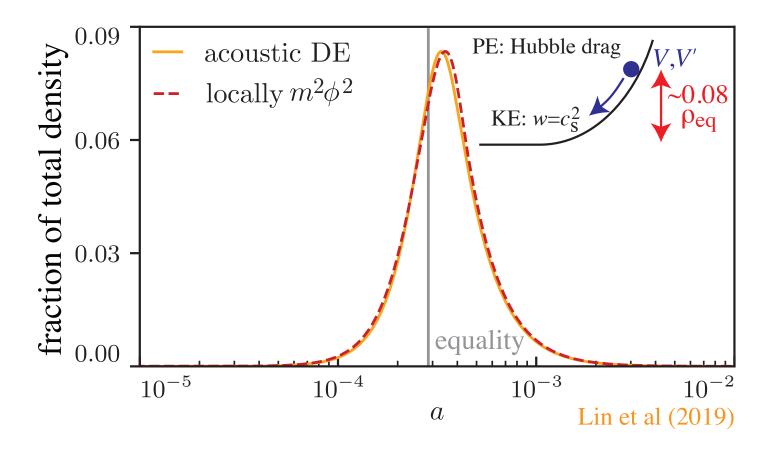
 Dark exotica with relativistic sound speed, acoustic oscillations, enhances decay



 Tune the impact with sound speed (first dark acoustic peak) and equation of state (redshifts away faster)

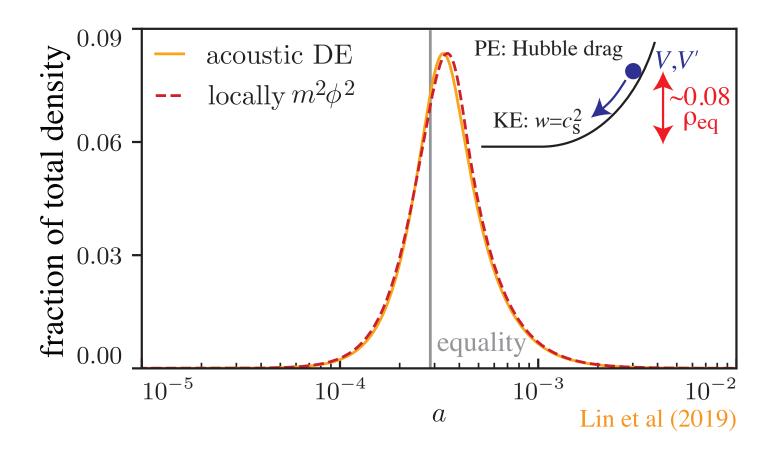


- Poulin et al (2018) 4 parameters, amplitude, time scale, c_s^2 and w must be carefully tuned (top of oscillatory potential $c_s^2 \downarrow$)
- Data favor something both more specific and generic: $c_s^2 = w$, transition to kinetic energy domination (Lin et al 2019)



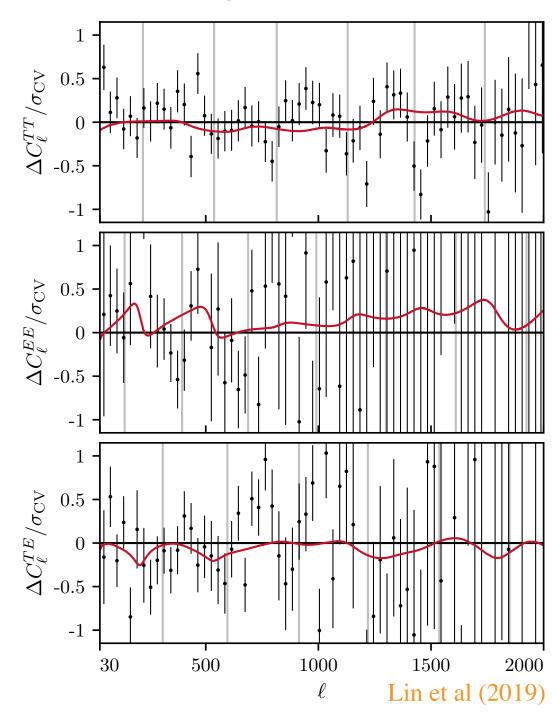
Potential to Kinetic

- With $c_s^2 = w = 1$ leaves 2 parameters: amplitude and slope of potential and requires kinetic energy redshift away (not oscillate)
- Amplitude $\sim 0.08 \rho_{\rm eq}$ and slope must be large enough to release from Hubble drag



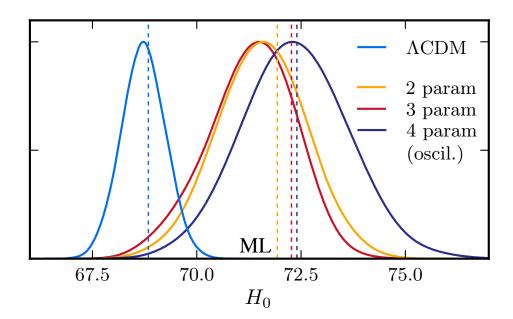
Fit and Predictive Signatures

- Fits joint data better by $\Delta \chi^2 \sim 12-14$ for 2-3 parameters
- Fits CMB itself
 better, largely TE
- TE glitch $\ell \sim 165$ highly sensitive
- Dark component redshifts away by recombination leaving nearly bare $\Omega_c h^2$ signature



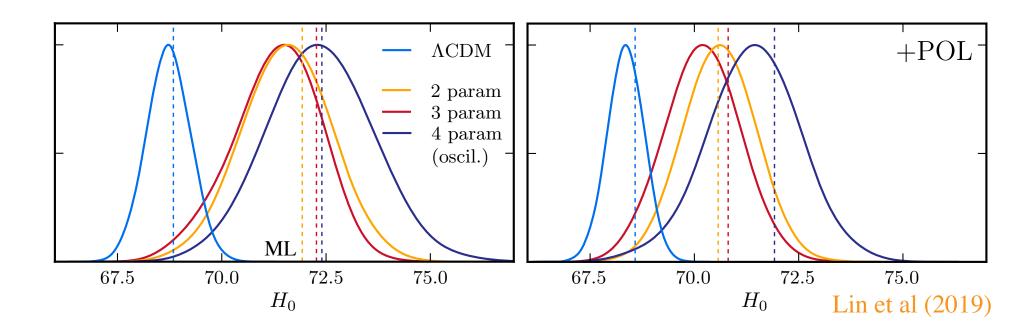
Potential Conversion of H_0 Tension

- Raises H_0 to bring CMB and local anchors into better agreement
- Minor further improvements with additional parameters
- Posterior depends on parameter volume near ΛCDM, maximum likelihood (ML) more reflective
- But mainly converts H_0 question to "why this, why then"!



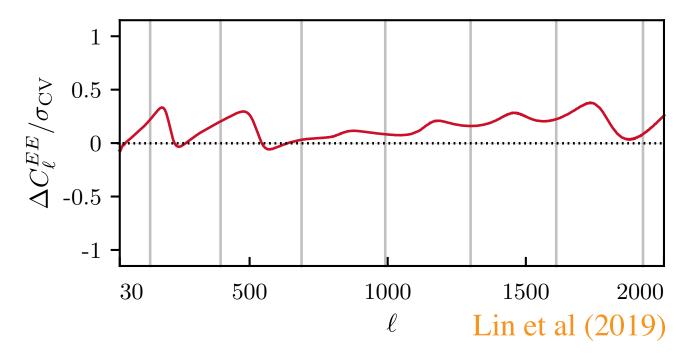
Potential Conversion of H_0 Tension

- Raises H_0 to bring CMB and local anchors into better agreement
- Minor further improvements with additional parameters
- Posterior depends on parameter volume near ΛCDM, maximum likelihood (ML) more reflective
- Already limited by Planck TE polarization, distinguishing details



Potential Conversion of H_0 Tension

- Raises H_0 to bring CMB and local anchors into better agreement
- Minor further improvements with additional parameters
- Posterior depends on parameter volume near ΛCDM, maximum likelihood (ML) more reflective
- EE residuals are ~ 0.3 vs cosmic variance per multipole



Opportunity for testing ideas based on changing CMB anchor!

Summary

- Planck and other precision CMB experiments have firmly established Λ CDM as the standard model
- ACDM unreasonably effective and efficient in describing suite of cosmological observables
- 6 numbers, mostly measured to sub percent precision, mostly consistent at this level with everything
- Tensions, anomalies and curiosities: imperfection is more interesting than perfection
- H_0 at 4.4σ , can only be explained by changing one of the anchors
- CMB anchor is sound horizon and cross checked by damping scale
- Potential conversion illustrates designer difficulties, one or more parameters per effect:
 - look for predictive power of any explanation